

Self-force corrections to spin precession for eccentric orbits in Schwarzschild spacetime (again)

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Last time on Capra

Numerical Results for $\Delta\psi(e = 0.1)$

p	$\Delta\psi$	$\lim_{e \rightarrow 0} \Delta\psi$	$(\Delta\psi - \lim_{e \rightarrow 0} \Delta\psi)/e^2$
10	$-5.0374746183 \times 10^{-2}$	-5.06715×10^{-2}	2.96711×10^{-2}
20	$-4.105502714 \times 10^{-2}$	-4.12414×10^{-2}	1.8634×10^{-2}
30	$-2.97664644 \times 10^{-2}$	-2.99044×10^{-2}	1.3792×10^{-2}
40	$-2.309158147 \times 10^{-2}$	-2.32001×10^{-2}	1.085×10^{-2}
50	$-1.880823903 \times 10^{-2}$	-1.88976×10^{-2}	8.93×10^{-3}
60	$-1.58482057 \times 10^{-2}$	-1.59242×10^{-2}	7.595×10^{-3}
70	$-1.3686455636 \times 10^{-2}$	-1.37523×10^{-2}	6.588×10^{-3}
80	$-1.204035626 \times 10^{-2}$	-1.20966×10^{-2}	5.826×10^{-3}
90	$-1.074624358 \times 10^{-2}$	-1.07983×10^{-2}	5.204×10^{-3}
100	$-9.702092699 \times 10^{-3}$	-9.74941×10^{-3}	4.732×10^{-3}
110	$-8.84229541 \times 10^{-3}$	-8.88571×10^{-3}	4.341×10^{-3}
120	$-8.122988585 \times 10^{-3}$	-8.16224×10^{-3}	3.925×10^{-3}
130	$-7.511573415 \times 10^{-3}$	-7.54748×10^{-3}	3.59×10^{-3}
140	$-6.985103217 \times 10^{-3}$	-7.01869×10^{-3}	3.359×10^{-3}
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WRONG!

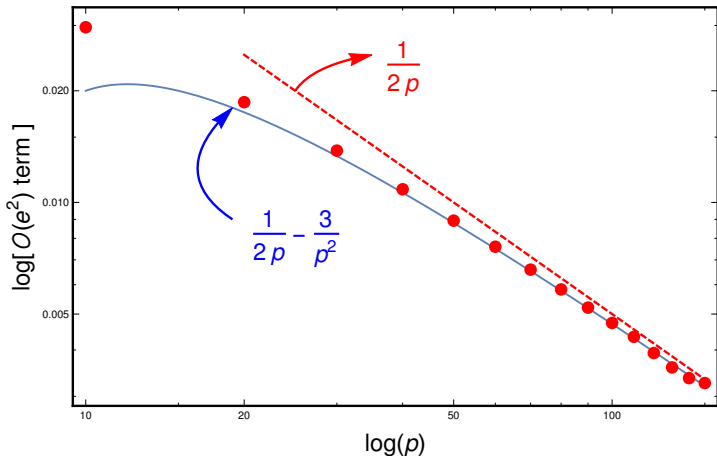
Last time on Capra

and
$$\Delta\psi - \lim_{e \rightarrow 0} \Delta\psi = \frac{1}{2} \frac{e^2}{p}$$

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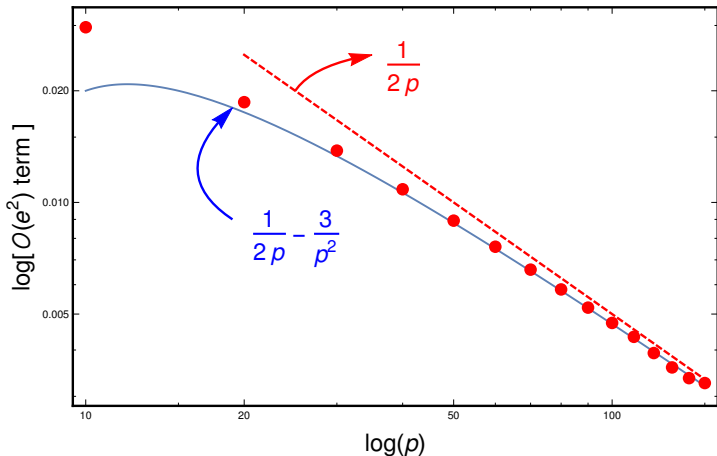
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Last time on Capra

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$$\Delta\psi - \lim_{e \rightarrow 0} \Delta\psi = \frac{1}{2} \frac{e^2}{p} \text{ vs. NO } 1/p \text{ term (PN L.O.)}$$



Recalling the formalism I

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ψ is the accumulated precession per angle between **two** frames.

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- 1 //-transported tetrad: $\dot{\lambda}_i^\alpha = 0$ $\{u^\alpha, \lambda_1^\alpha, \lambda_2^\alpha \propto \hat{\theta}, \lambda_3^\alpha\}$
- 2 Marck's tetrad (Marck 1983): $\{u^\alpha, e_1^\alpha, e_2^\alpha \propto \hat{\theta}, e_3^\alpha\}$

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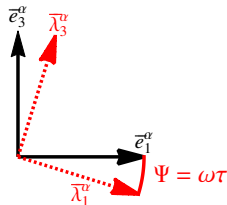
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Rotation in the 1 – 3 plane by Ψ

$$\begin{pmatrix} \lambda_1^\alpha \\ \lambda_3^\alpha \end{pmatrix} = \begin{pmatrix} \cos \Psi & -\sin \Psi \\ \sin \Psi & \cos \Psi \end{pmatrix} \begin{pmatrix} e_1^\alpha \\ e_3^\alpha \end{pmatrix}$$

Using $\dot{\lambda}_1 = \dot{\lambda}_3 = 0$ we get $\dot{\Psi} = \boxed{e_{3\alpha} \dot{e}_1^\alpha = -e_{1\alpha} \dot{e}_3^\alpha}$



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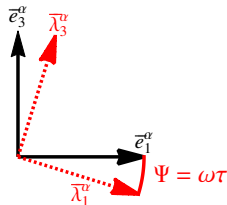
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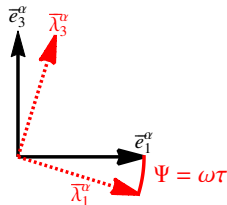
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Thus

$$\delta \dot{\Psi} = \delta \left(g_{\alpha\beta} e_3^\beta u^\gamma \nabla_\gamma e_1^\alpha \right) = -\delta \left(g_{\alpha\beta} e_1^\beta u^\gamma \nabla_\gamma e_3^\alpha \right)$$

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$$\left. \begin{aligned} \delta\Psi &= \int_0^{2\pi} \frac{d\bar{\Psi}}{d\chi} \left(\frac{\delta\dot{\Psi}}{\bar{\Psi}} - \frac{\delta\dot{R}}{\bar{u}^r} \right) d\chi \\ \delta\Phi &= \int_0^{2\pi} \frac{d\bar{\Phi}}{d\chi} \left(\frac{\delta L}{\bar{L}_0} - \frac{\delta\dot{R}}{\bar{u}^r} \right) d\chi \end{aligned} \right\}$$

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Progress since Capra 2015: missing terms

The **crux**: numerical computation of

$$\delta\dot{\Psi} = \underbrace{\delta\left(g_{\alpha\beta}e_3^\beta u^\gamma \nabla_\gamma e_1^\alpha\right)}_{\delta\dot{\Psi}_1} = -\underbrace{\delta\left(g_{\alpha\beta}e_1^\beta u^\gamma \nabla_\gamma e_3^\alpha\right)}_{\delta\dot{\Psi}_2},$$

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where

$$\delta\dot{\Psi}_1 = \frac{1}{2} \dot{\Psi} (h_{uu} + h_{33} - h_{11}) + c_{01} \bar{\Gamma}_{311} + c_{03} \bar{\Gamma}_{313} + \delta\Gamma_{310} + \frac{dc_{13}}{d\tau},$$

$$\delta\dot{\Psi}_2 = \frac{1}{2} \dot{\Psi} (h_{uu} + h_{11} - h_{33}) - c_{01} \bar{\Gamma}_{131} - c_{03} \bar{\Gamma}_{133} - \delta\Gamma_{130} - \frac{dc_{31}}{d\tau},$$

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$$\text{E.g. } c_{03}(\chi) = \text{func.}(p, e, \chi) \times \int_0^\chi \frac{d\tau}{d\chi'} F_\phi^{\text{cons}}(\chi') d\chi'$$

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\Rightarrow Integrand is zero or integrates to zero.

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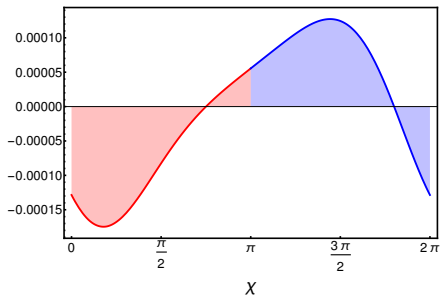
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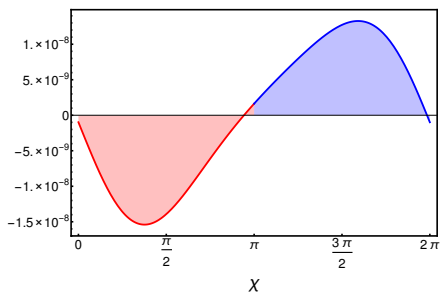
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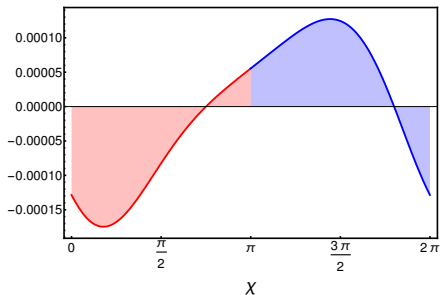
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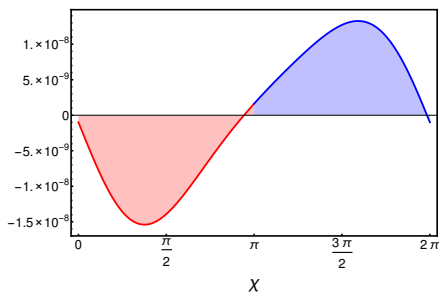
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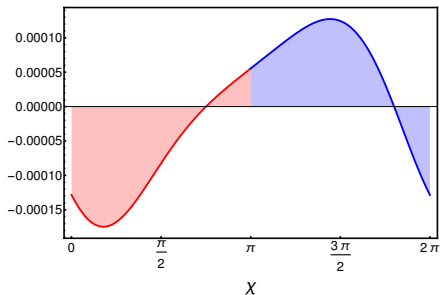
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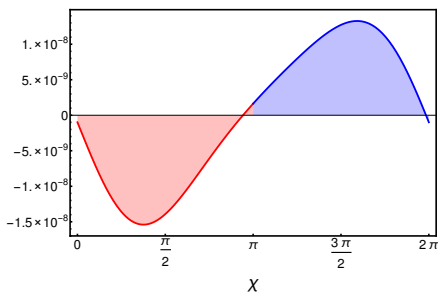
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$$\delta\Psi_1 - \delta\Psi_2 \lesssim \mathcal{O}(10^{-10}) \text{ for } p \in [10, 100], e \leq 0.2$$

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Data for $p = \{10, 20, 30, \dots, 100\}$, $e = \{0.05, 0.1, 0.15, 0.2\}$

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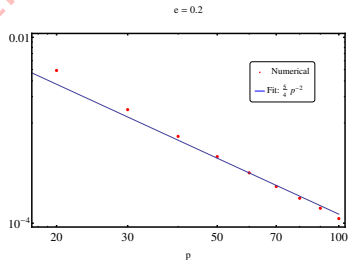
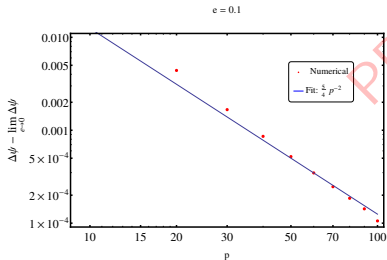
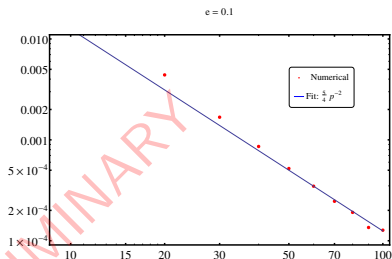
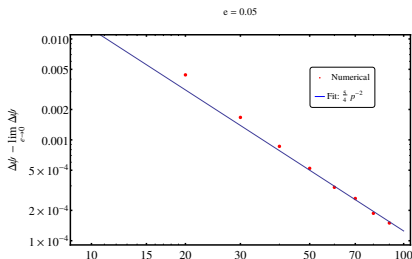
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Difference is **gauge invariant** and its physical origin understood (fixing **two** frequencies).

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Extract the $\mathcal{O}(e^2)$ term $\sim \frac{5}{4} \frac{e^2}{p^2}$. Compare with future **1PN** result.



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Consistent with current PN knowledge
i.e. NO p^{-1} terms at $\mathcal{O}(e^2)$.

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- **Eccentric, tidal** gauge invariants.